

Venus Express en route to probe the planet's hidden mysteries

The European spacecraft Venus Express has been successfully placed into a trajectory that will take it on its journey from Earth towards its destination of the planet Venus, which it will reach next April.

A virtual twin sister of the Mars Express spacecraft which has been orbiting the Red Planet since December 2003, Venus Express is the second planet-bound probe to be launched by the European Space Agency.

Venus Express will eventually manoeuvre itself into orbit around Venus in order to perform a detailed study of the structure, chemistry and dynamics of the planet's atmosphere, which is characterised by extremely high temperatures, very high atmospheric pressure, a huge 'greenhouse effect' and as-yet inexplicable 'super-rotation' which means that it speeds around the planet in just four days.

The European spacecraft will also be the first orbiter to probe the planet's surface while exploiting the 'visibility windows' recently discovered in the infrared waveband.

The 1240 kg mass spacecraft was developed for ESA by a European industrial team led by EADS Astrium with 25 main contractors spread across 14 countries. It lifted off on board a Soyuz- Fregat rocket, the launch service being provided by Starsem.

Venus Express is currently distancing itself from Earth at full speed, heading on its fivemonth, 350 million kilometre journey inside our Solar System. After check-outs to ensure that its on-board equipment and instrument payload are in proper working order, the spacecraft will be 'mothballed', with contact with Earth being reduced to once daily. If needed, trajectory correction manoeuvres can go ahead at the half-way stage in January.

When making its closest approach, Venus Express will face far tougher conditions than those encountered by Mars Express on nearing

the Red Planet. For while Venus's size is indeed similar to that of Earth, its mass is 7.6 times that of Mars, with gravitational attraction to match.

To resist this greater gravitational pull, the spacecraft will have to ignite its main engine for 53 minutes in order to achieve 1.3 km/second deceleration and place itself into a highly elliptical orbit around the planet. Most of its 570 kg of propellant will be used for this manoeuvre.

A second engine firing will be necessary in order to reach final operational orbit: a polar elliptical orbit with 12-hour crossings. This will enable the probe to make approaches to within 250 km of the planet's surface and withdraw to distances of up to 66 000 km, so as to carry out close-up observations and also get an overall perspective.

The Venus Express spacecraft Exploring other planets to better understand planet Earth.

The Venus Express mission is planned to last at least two Venusian days (486 Earth days) and may be extended, depending on the spacecraft's operational state of health.

Venus Express largely re-uses the architecture developed for Mars Express. This has reduced manufacturing cycles and halved the mission cost, while still targeting the same scientific goals. Finally approved in late 2002, Venus Express was thereby developed fast, indeed in record time, to be ready for its 2005 launch window.

However, Venusian environmental conditions are very different to those encountered around Mars. Solar flux is four times higher and it has been necessary to adapt the spacecraft design to this hotter environment, notably by entirely redesigning the thermal insulation.

Whereas Mars Express sought to retain heat to enable its electronics to function properly, Venus Express will in contrast be aiming for maximum heat dissipation in order to stay

The solar arrays on Venus Express have been completely redesigned. They are shorter and are interspersed with aluminium strips to help reject some solar flux to protect the spacecraft from temperatures topping 250°C.

It has even been necessary to protect the rear of the solar arrays — which normally remain in shadow — in order to counter heat from solar radiation reflected by the planet's atmosphere.

Following on from the twenty or so American and Soviet missions to the planet carried out since 1962, Venus Express will endeavour to answer many of the questions raised by previous missions but so far left unanswered. It will focus on the characteristics of the atmosphere, its circulation, structure and composition in relation to altitude, and its interactions with the planet's surface and with the solar wind at altitude.

PERTH OBSERVATORY VOLUNTEERS' GROUP

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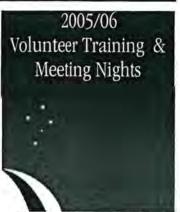
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Below are listed the training nights scheduled for 2006.

The vollies decided at a recent meeting that they would like to rationalise the training/meeting night schedule. They decided that scheduling a meeting on the third Monday of the month would be more convenient. We have checked for clashes with Observatory events and so far none have been identified. (Naturally, flexibility needs to be maintained and the meeting date will be changed from the strict schedule as the need arises.)

Training is important for our volunteers, they enjoy it, and we need to support these staff members in return for the assistance they render.

All staff are directed NOT to ROSTER
ANY ACTIVITY THAT WILL INTERFERE
WITH TRAINING night attendees access to
the lecture theatre and the VOF (and associated telescopes).

Also, could you please bring the details of this message to the attention of all staff under your direct supervision.

Your co-operation is appreciated.

Jamie Biggs.

16th January (Changed from 23rd January) 20th February (Changed from 27th Feb)

20th March (Changed from 27th March)

18th April (Note: Tuesday)

15th May (Changed from 15th May)

19th June

17th July

21st August

18th September

16th October

20th November

18th December (Maybe we'll cancel this one - it's very close to Xmas)

Stardust returns home

After a 2.88-billion-mile (4.63 billion kilometers) journey that would astonish the most dependable postal worker, NASA's Stardust spacecraft completed a delivery that was years in the making. The payoff? Scientists now have the first direct samples of cometary material.

An infrared camera captures Stardust's sample-return capsule as it is descends to the Utah Test and Training Range. This image was taken from NASA TV.

NASA [larger image]

At 2:12 a.m. PST, Stardust released the sample-return capsule, which weighs about 100 pounds (45.36 kilograms). The container entered Earth's atmosphere above the Pacific Ocean at nearly 29,000 mph (46,000 kilometers per hour) — the fastest reentry of a probe ever. The capsule slowed and released a drogue parachute about 20 miles (32 km) above the surface. The main chute deployed after the vessel descended to 1.9 miles (3 km) above Utah. The capsule decelerated and safely landed at the U.S. Air Force's Utah Test and Training Range, southwest of Salt Lake City.

Recovery helicopter Vertigo 1 first reached the landing site and searched the area. Locating the vessel was difficult because of the darkness and the capsule, about the size of an automobile tire, is the same color as the salty desert floor. Eventually, the container was located, dou-

ble bagged, placed aboard the helicopter, and taken to a clean room at the nearby Dugway Proving Ground. NASA officals commented that the reentry, descent, and recovery occurred flawlessly.

In January 2004, Stardust passed through the coma of Comet 81P/Wild (pronounced "Vilt") 2. Using its tennis-racket-shape collector, the probe gathered approximately an ounce (28 grams) of dust grains from the comet's coma. Aerogel, a thin substance attached to the collector, captured the particles as they sped by at 5 times the speed of a rifle bullet. The comet samples were then secured in a return capsule, which was safely stowed until today.

Stardust contributed greatly to astronomers' knowledge of comets, long before the sample return. In 2004, it returned close-up views, including the nucleus, of Wild 2. The results amazed scientists.

"We thought Comet Wild 2 would be like a dirty, black, fluffy snowball," explains Stardust Principal Investigator Donald Brownlee. "Instead, it was mind-boggling to see the diverse landscape in the first pictures from Stardust, including spires, pits, and craters, which must be supported by a cohesive surface."

Stardust also flew within about 2,050 miles (3,300 km) of the main-belt asteroid Annefrank before it met Comet Wild 2. The

observations revealed the asteroid is darker than expected and about 5 miles (8 km) long — approximately twice the size predicted from Earth-based observations.

Why send a spacecraft to meet a comet? Comets are relics from when the solar system formed around 4.5 billion years ago. By learning more about comets, scientists will understand how the solar system formed and evolved.

Although the capsule is safe on Earth, the journey is not over for Stardust. About 20 minutes after it jettisoned the container, the spacecraft fired its thrusters and moved into a solar orbit. At this time, NASA has no plans to extend Stardust's mission.

Beyond thrilling scientists with the delivery of comet dust, Stardust also delighted many skygazers in the western United States. As the return capsule darted across the sky shortly before 2 a.m. PST, people living under clear skies in an area from San Francisco to Salt Lake City to Seattle had the chance to see it. The capsule reached maximum heating and became its brightest as it passed over Nevada. It took the capsule only about a minute to travel from the northwest coast of California to the landing site in northwestern Utah.

SOURCE:http://www.astronomy.com/asy/def ault.aspx?c=a&id=3848

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Vollie Group Minutes

Perth Observatory Volunteer Group Inc

Minutes of Meeting November 28th 2005-11-29
Present. G.Lowe. E.Walker. J.Morris. L.Martin. M.Freeman.
N.Townsend. B.Harris.

J.Biggs. D.Emrich. A.MacNaughton. T.Beston.

Apologies, J.Alcroft D.Alderson, G.Coletti M.Zengerer, R.Bolen, FBilki

Confirmation of Minutes. The Minutes were confirmed as a true and correct record

Moved B.Harris seconded L.Martin.

Business Arising from the Minutes.

 M.Freeman referred to the proposed Corporate Viewing Night and requested further information. N.Townsend stated that the night was scheduled for December 8th at 8pm, Invitations had been sent to a number of well known personalities, and todate there had been 13 acceptances. J.Biggs outlined the format for the evening and hoped that a number of Volunteers would be present, and requested that as the evening was an up-market presentation, everyone attending should be neatly dressed, with Volunteers using their name badges.

2. Car Stickers . L. Martin stated that he could get a limited number

of stickers

printed, and requested details on colour and wording, after discussion it was agreed that the wording should be "GET STARRY EYED. VISIT PERTH OBSERVATORY" and if possible to incorporate the telephone number,

colour should be blue or black on a white background.

Treasurers Report, B.Harris said that there was no change in the financial situation.

General Business. J.Biggs gave details of the launch of the Internet Telescope Project which would occur on December 7th at 8.30pm with Minister Judy Edwards attending, all Volunteers and their families were welcome to attend.

N.Townsend tabled a number of examples of promotional material she had been able to get placed in various local papers, J.Biggs asked that all items in the Press regarding the Observatory should be saved, and to ensure that the Observatory had a copy

There being no further General Business the Meeting closed at 7.40pm

Next Meeting will be on January 16th 2006

The Splendor of Orion: A Star Factory Unveiled

Scientists have created the most detailed portrait ever of the closest known star factory, the Orion Nebula. They have also uncovered new details about the stellar winds responsible for carving out the nebula's ghostly shapes.

Meanwhile, another research group identified a glowing gas cloud in another region of the night sky that might replace Orion when it fades from view in about 100,000 years.

Orion is located 1,500 light-years away and is the nearest region of massive star-formation to Earth, Located at the center of the Orion Nebula is a group of four young, massive stars; they are collectively referred to as the "Trapezium" because of the shape they make. The Trapezium is surrounded by a halo of 1,000 faint, low-mass stars similar to the Sun.

Like all massive stars, those in the Trapezium are radiating a stream of energetic X-rays called a stellar wind. However, the wind from the Trapezium is millions of times denser and more energetic than the flow from our own Sun.

When these powerful winds collide with the dense cloud of dust and gas surrounding the stars, they produce shock waves that erode and shape the clouds. Similar to how winds on Earth can sculpt sand dunes in a desert, stellar winds from the Trapezium are carving out a relatively dust and gas free cavity filled with strange, ethereal scenes.

"Orion may seem very peaceful on a cold winter night, but in reality it holds very massive, luminous stars that are destroying the dusty gas cloud from which they formed," said astronomer Tom Megeath.

The Trapezium's stellar winds play a major role in the creation of new stars within the Orion Nebula. If the winds are strong enough, their shockwaves act as catalysts for star formation by pushing the clouds together into compact bundles.

Stellar winds in two of Orion's three star forming regions and found that some of the winds have been blowing continuously for nearly 1,500 years.

The Trapezium's winds are doing such a thorough job of carving up its cloudy cocoon that in about 100,000 years, the Orion Nebula as we know it will fade from view, scientists say.

Stargazers of the future need not despair, however, because astronomers have discovered a gas cloud in the constellation Cassiopeia that has just begun shining with newborn stars. Known as W3, the cloud will become the Grand Nebula and replace Orion as Earth's most visible stellar nursery just as Orion disappears.

The so-called proto-Trapezium stars are much younger than those found in Orion and appear to still be growing by accreting gas from their surrounding clouds.

The discovery of these stars may resolve a debate among scientists about how Orion's Trapezium formed in the first place. The Orion Nebula and the Trapezium within are of special interest to astronomers because it is believed that our own Sun was born in such an environment.

According to one Trapezium formation hypothesis, each of the four massive stars formed separately but then descended into the center of the nebula to form a clump. Another leading hypothesis is that the Trapezium stars never moved at all and formed together right where they are.

Because the proto-Trapezium stars in W3 are clustered together, Megeath's team suspects that the second hypothesis is correct.

When they mature, the proto-Trapezium stars will emit their own, powerful stellar gusts and ultraviolet light and carve out their own spectral skyscapes. Scientists are uncertain when this process will be complete or when the stars will reach their maximum brilliance, but think that it may around the time that Orion is scheduled to disappear.